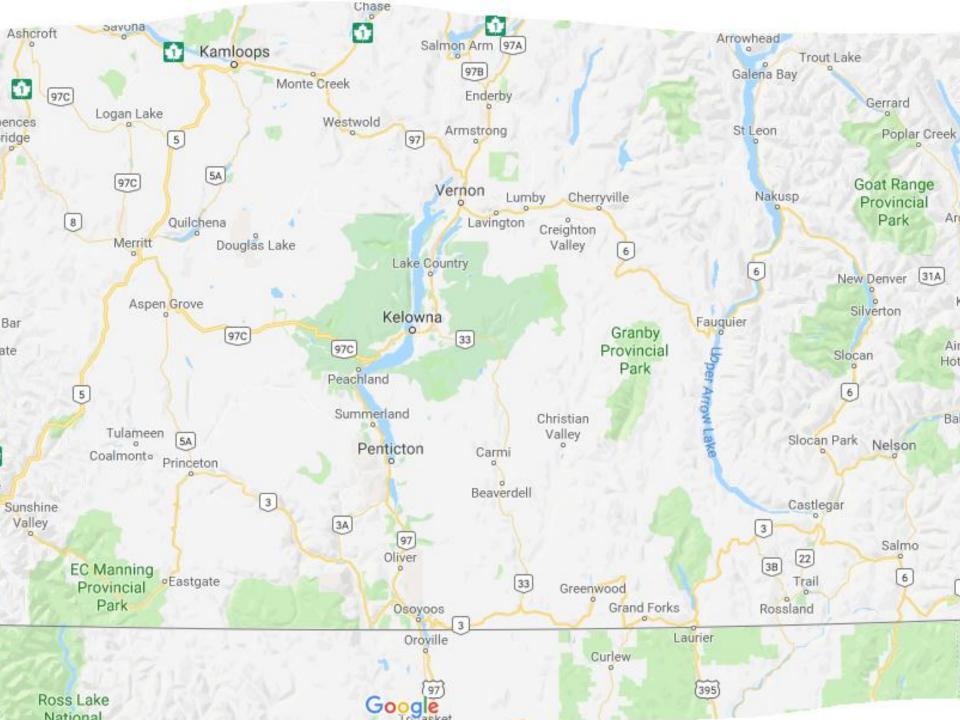
# The DRAO (Dominion Radio Astrophysical Ob.) Penticton, BC

**PAA Meeting** 

March 1<sup>st</sup>, 2019

By Astro-Traveler - Rick Stankiewicz



### **Canada's Observatories**

#### NAC-CNAC

**Herzberg Astronomy and Astrophysics** 











### Site Map / Plan du site

CHIME Pathfinder Telescope Télescope de CHIME Pathfinder

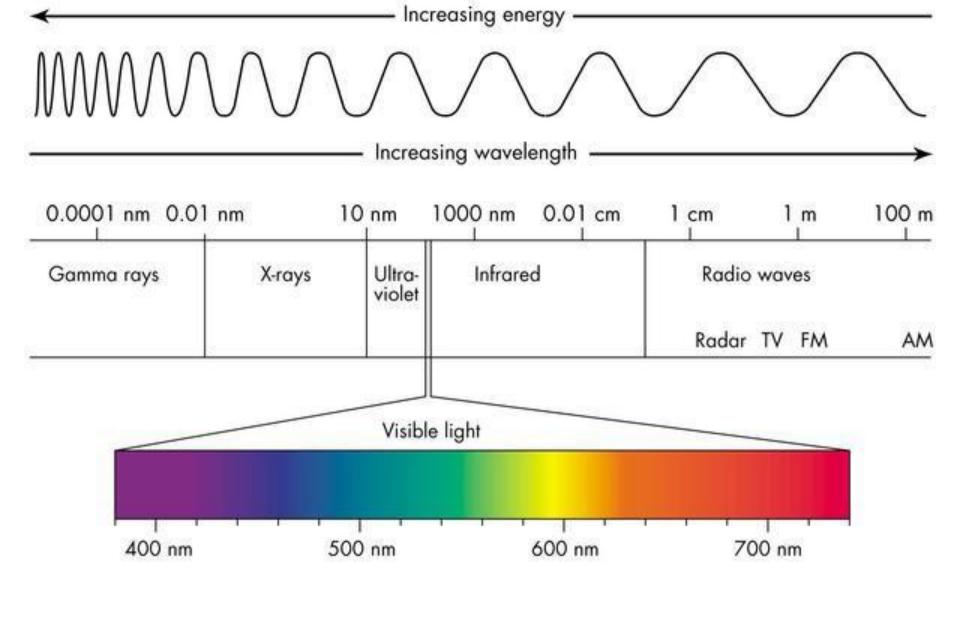






















### Solar Telescope

#### Solar Flux Monitor Program

The Solar Flux Monitor Program measures the strength of the radio waves produced by the Sun. It is the most accurate and longest running measure of the Sun's activity in existence. Astronomers worldwide use the results in their research. The data is of special interest to industries that are affected by the Sun – power and radio communications companies, and satellite operators. The program ...

- o measures the radio emission from the Sun at a wavelength of 10.7 cm
- o originated in 1946 near Ottawa and moved to DRAO in 1990
- o is a fully automated system that has tracked the sun for over 50 years

Le Programme de monitoring du flux radio solaire

Le Programme de monitoring du flux radio solaire mesure l'intensité des ondes radio produites par le soleil. Il s'agit de la mesure à plus long terme et la plus précise de l'activité solaire ayant jamais existé. Ses résultats sont utilisés à des fins de recherche par des astronomes du monde entier. De plus, les données recueillies sont d'un intérêt particulier pour les industries affectées par l'activité solaire, soit les compagnies énergétiques et de communication radio et les opérateurs de satellites. Ce programme :

- o permet de mesurer l'émission radio en provenance du soleil à une longueur d'onde de 10.7 cm
- o a débuté à Ottawa en 1946 et a été déménagé à l'OFR en 1990
- o est un système entièrement automatisé qui observe le soleil depuis plus de 50 ans.



le Programme de monitoring du flux radio solaire





### Synthesis Telescope



The Synthesis Telescope comprises of seven 9m dishes on an east-west line that is 600m long. Three of the antennas can be moved to different locations along a precision rail track.

This telescope began in 1972 as a two-element interferometer and was quickly expanded to four antennas. A series of upgrades between 1985 and 1994 led to an additional observing frequency, expansion of the system to seven antennas, and new correlators that were designed and built by DRAO engineers.

The Synthesis Telescope operates as a national facility for facility astronomy and has recently been mostly dedicated to the Canadian Galactic Plane Survey (CGPS).





# TÉLESCOPE JOHN A. GALT TELESCOPE



### 26m Telescope



Construction began on this telescope in 1959 and was completed in time for the official opening of DRAO in 1960. The dish is 25.6m in diameter, fully steerable, has an open-mesh surface, and can observe at multiple frequencies.

The 26m telescope was used in the first Very Long Baseline Interferometry (VLBI) experiment in 1967 in conjunction with the 46m telescope at Algonquin Radio Observatory in Ontario.

This telescope has been used over the decades to provide data to cuttingedge science projects, including the Canadian Galactic Plane Survey (CGPS).

Over the past 3 years the receiving system of the 26m telescope has been upgraded for the Global Magneto-Ionic Medium Survey (GMIMS).

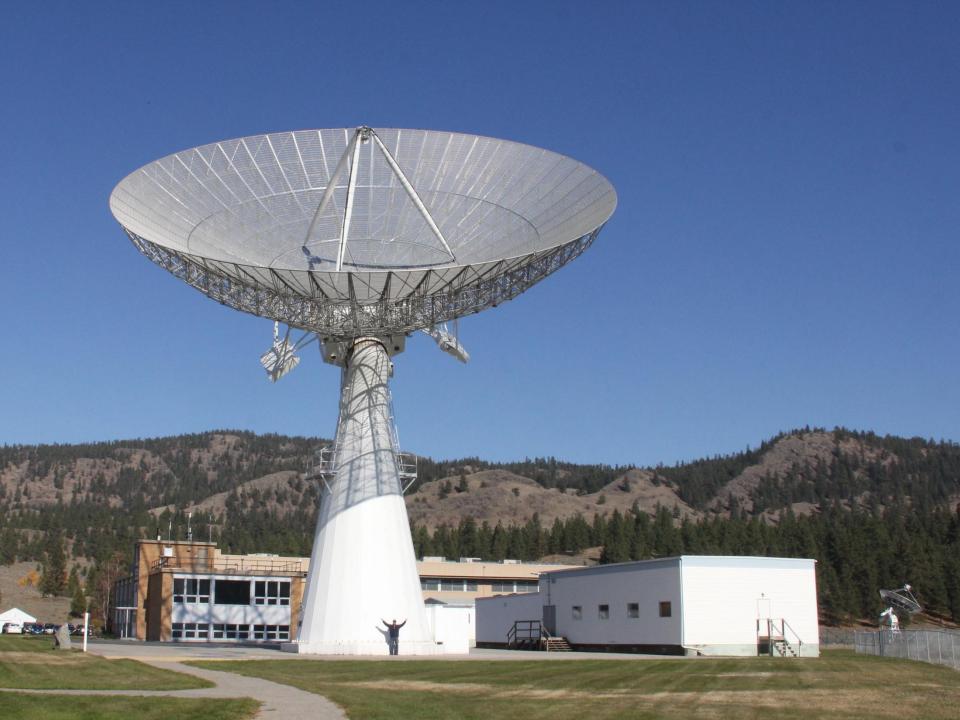
### IEEE MILESTONE IN ELECTRICAL ENGINEERING AND COMPUTING

First Radio Astronomical Observations Using VLBI, 1967

On the morning of 17 April 1967, radio astronomers used this radiotelescope at DRAO and a second one at the Algonquin Radio Observatory located 3074 km away to make the first successful radio astronomical observations using Very Long Baseline Interferometry. Today, VLBI networks span the globe, extend into space, and continue to make significant contributions to both radio astronomy and geodesy.

September 2010











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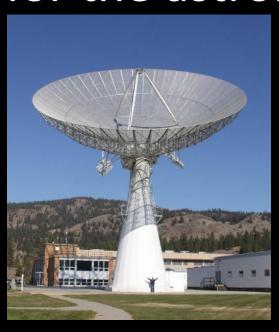






## Hope you enjoyed the tour of the DRAO

When you travel, make sure you plan for the astronomical and share with



the PAA



Keep looking up!